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Generalized Gibbs Ensemble of 2d CFTs at large central charge in the thermodynamic limit

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ABSTRACT: We discuss partition function of 2d CFTs decorated by higher qKdV charges in the thermodynamic limit when the size of the spatial circle goes to infinity. In this limit the saddle point approximation is exact and at infinite central charge generalized partition function can be calculated explicitly. We show that leading 1/c corrections to free energy can be reformulated as a sum over Young tableaux which we calculate for the first two qKdV charges. Next, we compare generalized ensemble with the "eigenstate ensemble" that consists of a single primary state. At infinite central charge the ensembles match at the level of expectation values of local operators for any values of qKdV fugacities. When the central charge is large but finite, for any values of the fugacities the aforementioned ensembles are distinguishable.

KEYWORDS: Conformal Field Models in String Theory, Conformal Field Theory, Conformal and W Symmetry, Holography and condensed matter physics (AdS/CMT)

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1 Introduction

Conformal Field Theories in two dimensions describe one dimensional systems at criticality, most notably string theoretic worldsheet. Thermalization of two-dimensional theories came to the focus of attention recently in the context of entanglement spreading following a quantum quench [1–3]. Additional motivation to study 2d theories comes from holographic duality, which relates thermal properties of boundary CFTs to black hole physics in AdS₃ [4–14]. Thermal partition function of 2d CFTs on a torus is a well-studied object which exhibits elegant mathematical properties making a connection with the theory of modular forms [15, 16]. In the thermodynamic limit, when the size of spatial circle ℓ goes to infinity, torus degenerates into a cylinder and thermal partition function becomes particularly simple. At the level of free energy it only depends on central charge c and inverse temperature β ,

$$F = \frac{c\,\pi^2\ell}{6\,\beta}.\tag{1.1}$$

The peculiar property of conformal field theories in two dimensions is that, in addition to energy, there is an infinite number of local conserved charges associated with quantum KdV algebra [17–19]. Whenever a system has additional conserved quantity, e.g. total

number of particles, its thermal properties are described by the grand canonical ensemble, with the effective number of particles controlled by the value of chemical potential [20]. In principle similar logic applies when there are many conserved quantities. When the system is integrable the number of conserved quantities becomes infinite, raising the question if such a system would even approach thermal equilibrium, and what that equilibrium state might be. These questions have drawn significant attention recently because of experimental realizations of such a scenario [21]. The prevailing view is that integrable system would evolve toward a state described by the Generalized Gibbs Ensemble (GGE) [22–24], which includes all conserved mutually-commuting charges Q_k ,

$$Z = \operatorname{Tr}\left(e^{-\sum_{k} \mu_{k} Q_{k}}\right), \qquad Q_{1} = H, \qquad \mu_{1} = \beta.$$
(1.2)

The value of corresponding chemical potentials, also called fugacities, μ_k should be chosen such that the ensemble (1.2) would have the same value of conserved charges as the initial state. Applying this logic to the two-dimensional CFTs it is natural to expect that e.g. following a quantum quench these theories would thermalize to a state which has the same local properties as the GGE decorated by higher qKdV charges [25–31]. This expectation motivates our study of GGE of two-dimensional theories in this paper.

We start with a CFT on a spatial circle of length ℓ , such that Euclidean space-time is parametrized by a holomorphic coordinate $\omega = u + it$. The latter is related to the conventional coordinate on the plane via $z = e^{\frac{2\pi i}{\ell}\omega}$. In addition to the Hamiltonian

$$H = Q_1 = \int_0^\ell du \, T(u), \tag{1.3}$$

conformal theory always has an infinite tower of mutually commuting qKdV charges [17]

$$Q_3 = \int_0^\ell du : T(u)^2 :, \qquad Q_5 = \int_0^\ell du : T(u)^3 : + \frac{c+2}{12} : T'(u)^2 :, \qquad \dots$$
 (1.4)

By convention qKdV charges are parametrized by odd numbers Q_{2k-1} , $k \geq 1$. We are interested in calculating

$$Z = \text{Tr}\left(e^{-\beta H - \sum_{k=2}^{\infty} \mu_{2k-1} Q_{2k-1}}\right). \tag{1.5}$$

When all $\mu_{2k-1} = 0$ for $k \geq 2$, (1.5) is the regular partition function on a torus with the modular parameter $\tau = i \frac{\beta}{2\pi\ell}$. In particular partition function is invariant under modular transformations of τ . Turning on μ_{2k-1} would break modular invariance, although Z still exhibits some interesting properties under modular transformations [32]. In this work we focus on thermodynamic limit $\ell \to \infty$, while β, μ_{2k-1} are kept fixed. Our goal is to calculate the extensive part of free energy $F = \log Z$,

$$F = \frac{c \pi^2 \ell}{6 \beta} f(\beta, \mu_{2k-1}, c) + o(\ell).$$
 (1.6)

We show that at infinite central charge f can be expressed in terms of an algebraic relation, while leading 1/c corrections can be reformulated as a straightforward sum over Young tableaux which leads to a concise and explicit answer (2.52).

In the second part of the paper we compare local properties of the GGE (1.5) with those of an individual energy eigenstate. We find that at infinite central charge local properties of a highly excited Virasoro primary eigenstates exactly match those of the GGE for any values of the fugacities μ_{2k-1} . In this sense heavy primary states can be regarded as thermal at leading order in c. At the same time when central charge is finite, GGE and a primary eigenstate are always distinguishable in terms of expectation values of local operators. In other words, it is impossible to tune the fugacities μ_{2k-1} such that the GGE would have the same values of qKdV charges Q_{2k-1} as a heavy CFT primary. This conclusion is intriguing because it clearly shows that local properties of 2d CFTs upon equilibration are not always governed by the generalized ensemble of the form (1.5). Our finding emphasizes that integrable structure of 2d theories may preclude them to thermalize in the sense that local properties of the equilibrated state may lack universality.

2 Generalized Gibbs Ensemble in the thermodynamic limit

The crucial simplification of thermodynamic limit $\ell \to \infty$ is that saddle point approximation becomes exact. We first illustrate that in the case of the conventional Gibbs ensemble

$$Z(\beta) = \text{Tr}(e^{-\beta H}), \quad H = \frac{L_0 - c/24}{\ell}.$$
 (2.1)

Here L_n is the conventional Virasoro algebra generator related to the stress tensor on the plane,

$$T(z) = \sum_{n} \frac{L_n}{z^{n+2}},\tag{2.2}$$

while the Hamiltonian of CFT on a circle is given by (1.3).¹ The sum in (2.1) goes over all states in the Hilbert space,

$$|m_i, \Delta\rangle = L_{-m_1} \dots L_{-m_k} |\Delta\rangle, \qquad m_i \ge m_j \text{ for } i \ge j,$$
 (2.3)

parametrized by the dimension of Virasoro primary Δ and sets $\{m_i\}$, $\sum_{i=1}^k m_i = n$, arranged in the dominance order. Using the degeneracy of H, $L_0|m_i,\Delta\rangle = (\Delta+n)|m_i,\Delta\rangle$ the partition function can be represented as a double sum,

$$Z(\beta) = \sum_{\Lambda} \sum_{n=0} P(n) e^{-\frac{\beta}{\ell}(\Delta+n)}.$$
 (2.4)

In what follows we are only interested in the extensive part of free energy, and therefore we dropped the explicitly c-dependent term in H. The sum \sum_{Δ} goes over all Virasoro primaries including possible multiplicities, and P(n) is the number of integer partitions — Young tableaux consisting of n elements. For large n [33],

$$P(n) \approx e^{\pi\sqrt{\frac{2n}{3}}} \tag{2.5}$$

¹The stress tensor on the cylinder T(u) is related to the stress tensor on the plane T(z) via the standard conformal transformation.

and the sum over n can be substituted by an integral. Similarly, the sum over Δ in (2.4) can be substituted by an integral multiplied by the density of primaries given by Cardy formula [34]. Combining all together gives

$$Z(\beta) = \int d\Delta \, e^{\pi \sqrt{\frac{2}{3}(c-1)\Delta}} \int dn \, e^{\pi \sqrt{\frac{2n}{3}}} e^{-\frac{\beta}{\ell}(\Delta+n)} = e^F, \quad F = \frac{c \, \pi^2 \, \ell}{6 \, \beta}. \tag{2.6}$$

It is easy to see that in the limit $\ell \to \infty$ the saddle point approximation is exact,

$$\mathcal{L} = \pi \sqrt{\frac{2}{3}(c-1)\Delta} + \pi \sqrt{\frac{2n}{3}} - \frac{\beta}{\ell}(\Delta + n), \tag{2.7}$$

$$\Delta^* = (c-1)\pi^2 \ell^2 / 6\beta^2, \quad n^* = \pi^2 \ell^2 / 6\beta^2, \qquad \frac{\partial \mathcal{L}}{\partial \Delta} \Big|_{\Delta^* = n^*} = \frac{\partial \mathcal{L}}{\partial n} \Big|_{\Delta^* = n^*} = 0, \tag{2.8}$$

$$F = \mathcal{L}(\Delta^*, n^*) = \frac{c \pi^2 \ell}{6\beta}.$$
 (2.9)

2.1 GGE at infinite central charge

Our next step is to decorate the partition function by higher qKdV charges,

$$Z(\beta, \mu_{2i-1}) = \text{Tr}\left(e^{-\beta H - \sum_{k} \mu_{2k-1} Q_{2k-1}}\right). \tag{2.10}$$

We start our analysis with Q_3 (1.4). The expression for Q_3 in terms of Virasoro generators can be found in [17] (in our case expression for Q_3 is different by an overall factor $1/\ell^3$). Using the explicit form of Q_3 we split it into two parts as follows

$$Q_3 = \hat{Q}_3 + \tilde{Q}_3 \tag{2.11}$$

$$\ell^3 \hat{Q}_3 = \left(L_0^2 - \frac{c+2}{12} L_0 + \frac{c(5c+22)}{2880} \right), \quad \ell^3 \tilde{Q}_3 = 2 \sum_{n=1}^{\infty} L_{-n} L_n.$$
 (2.12)

The term \tilde{Q}_3 is defined such that it annihilates primary states, $\tilde{Q}_3|\Delta\rangle = 0$, while the rest is degenerate and depends only on the combination $\Delta + n$. Using scaling with ℓ of the saddle point values $\Delta^* \sim \ell^2$ and $n^* \sim \ell^2$ we immediately find that L_0 -independent term from \hat{Q}_3 would give $1/\ell^3$ contribution to free energy, while $-\frac{c+2}{12}L_0$ will contribute as $\sim 1/\ell$. Hence in the thermodynamic limit these terms can be neglected. Assuming for simplicity that only μ_3 is turned on we get,

$$Z(\beta, \mu_3) = \sum_{\Lambda} \sum_{n=0} P(n) e^{-\frac{\beta}{\ell}(\Delta+n) - \frac{\mu_3}{\ell^3}(\Delta+n)^2 + \tilde{\mathcal{L}}(c, \mu_3/\ell^3, \Delta, n)},$$
(2.13)

$$e^{\tilde{\mathcal{L}}(c,\mu_3/\ell^3,\Delta,n)} \equiv \frac{1}{P(n)} \sum_{\{m\}=n} \langle m_i, \Delta | e^{-\mu_3 \tilde{Q}_3} | m_i, \Delta \rangle.$$
 (2.14)

In (2.14) the sum is over sets $\{m_i\}$ satisfying $\sum_i m_i = n$, i.e. over the partitions of n.

A crucial simplification, which will be justified in the next section, is that Q_3 does not contribute to free energy at leading order in 1/c expansion. Hence at infinite central charge

one can simply take $\tilde{\mathcal{L}}$ to be zero,

$$Z(\beta, \mu_3) = \int d\Delta \int dn \, e^{\mathcal{L}}, \tag{2.15}$$

$$\mathcal{L} = \pi \sqrt{\frac{2}{3}(c-1)\Delta} + \pi \sqrt{\frac{2n}{3}} - \frac{\beta}{\ell}(\Delta + n) - \frac{\mu_3}{\ell^3}(\Delta + n)^2.$$
 (2.16)

Adding higher charges can be done in a similar way. Thus Q_5 , also given in [17], can be split into

$$Q_5 = \hat{Q}_5 + \tilde{Q}_5 \tag{2.17}$$

$$\ell^5 \hat{Q}_5 = \left(L_0^3 - \frac{c+4}{8} L_0^2 + \frac{(c+2)(3c+20)}{576} L_0 - \frac{c(3c+14)(7c+68)}{290304} \right), \tag{2.18}$$

$$\ell^{5}\tilde{Q}_{5} = \sum_{\substack{n_{1}+n_{2}+n_{3}=0\\n_{1},n_{2},n_{3}\neq 0}} : L_{n_{1}}L_{n_{2}}L_{n_{3}} : + \sum_{n=1}^{\infty} \left(\frac{c+11}{6}n^{2} - 1 - \frac{c}{4}\right)L_{-n}L_{n} + \frac{3}{2}\sum_{r=1}^{+\infty} L_{1-2r}L_{2r-1}.$$

Normal ordering in: $L_{n_1}L_{n_2}L_{n_3}$: means that the operators ordered to satisfy $n_3 \geq n_2 \geq n_1$. Again \tilde{Q}_5 is chosen to annihilate primary states, while \hat{Q}_5 is a function of L_0 . It is easy to see from ℓ -scaling of saddle point values of Δ , n that only L_0^3 term from \hat{Q}_5 contributes to the extensive part of free energy. Similarly to the case of Q_3 , we drop \tilde{Q}_5 at leading order in c. The rest is straightforward and can be generalized to all Q_{2k-1} . After integrating over n the partition function can be reduced to the integral over $E = \Delta + n$,

$$Z(\beta, \mu_{2k-1}) = \int dE \, e^{\mathcal{L}_E}, \quad \mathcal{L}_E = \pi \sqrt{\frac{2}{3} cE} - \frac{\beta}{\ell} E - \frac{\mu_3}{\ell^3} E^2 - \frac{\mu_5}{\ell^5} E^3 - \dots$$
 (2.19)

It is easy to see than that free energy will depend on the inverse temperature β and fugacities μ_i only through the combinations

$$F = \frac{c \pi^2 \ell}{6 \beta} f_0(t_{2k-1}), \qquad t_{2k-1} = \left(\frac{c \pi^2}{6 \beta^2}\right)^{k-1} \frac{\mu_{2k-1}}{\beta}. \tag{2.20}$$

Free energy F admits perturbative expansion in t_{2i-1} at any value of c, while expansion in μ_{2k-1} breaks down for large central charge. In principle f_0 is given by the algebraic equation specifying the saddle point E^* of (2.19),

$$\mathcal{L}_E = \frac{c \,\pi^2 \ell}{6 \,\beta} s_0, \qquad E^* = \frac{c \,\pi^2 \ell}{6 \,\beta} e^*, \qquad (2.21)$$

$$s_0 = 2\sqrt{e} - e - t_3 e^2 - t_5 e^3 - \dots,$$
 (2.22)

$$f_0 = s_0(e^*), \qquad \frac{\partial s_0}{\partial e}\Big|_{e^*} = 0.$$
 (2.23)

In practice perturbative expansion of f_0 is easier to recover iteratively. We notice that at leading order in c the expectation values of Q_{2k-1} are given by $\ell(E^*/\ell^2)^k$. In terms of the partition function (2.19) this can be rewritten as a differential equality

$$-\ell^{-1}\partial_{\mu_{2k-1}}\log Z = (-\ell^{-1}\partial_{\beta}\log Z)^{k},$$
(2.24)

or in terms of variables t_{2k-1} ,

$$(f_0 + 3t_3\partial_{t_3}f_0 + 5t_5\partial_{t_5}f_0 + \dots)^2 + \partial_{t_3}f_0 = 0,$$
(2.25)

$$(f_0 + 3t_3\partial_{t_3}f_0 + 5t_5\partial_{t_5}f_0 + \dots)^3 + \partial_{t_5}f_0 = 0,$$
(2.26)

. . .

For the Taylor expansion of f_0 equations (2.25), (2.26), etc. yield iterative relations which can be easily solved,

$$f_0 = 1 - t_3 + 4t_3^2 - 24t_3^3 - t_5 + 12t_3t_5 - 132t_3^2t_5 + 9t_5^2 - 234t_3t_5^2 - 135t_5^3 + \dots$$
 (2.27)

An elegant expression for f_0 in terms of the perturbative series in t_{2k-1} was given recently in [35].

2.2 GGE at finite central charge

When the central charge is finite, extensive part of free energy acquires 1/c corrections,

$$F = \frac{c \pi^2 \ell}{6 \beta} \left(f_0 + \frac{f_1}{c} + \frac{f_2}{c^2} + \dots \right). \tag{2.28}$$

Functions f_k admit Taylor expansion in terms of t_{2k-1} (2.20). For $k \geq 2$, f_k depends only on t_{2i-1} , $i \geq k$, and start with a term linear in t_{2k-1} . Thus, if expressed in terms of μ_{2k-1} , extensive part of free energy is polynomial in c. To calculate f_1, f_2, \ldots we need to take into account additional contribution to the effective action,

$$e^{\tilde{\mathcal{L}}(c,\mu_{2k-1}/\ell^{2k-1},\Delta,n)} \equiv \frac{1}{P(n)} \sum_{\{m\}=n} \langle m_i, \Delta | e^{-\mu_3 \tilde{Q}_3 - \mu_5 \tilde{Q}_5 - \dots} | m_i, \Delta \rangle.$$
 (2.29)

2.2.1 1/c corrections from \tilde{Q}_3

We start with the case when only $\mu_3 \neq 0$, (2.14). It is easy to see that the operator \tilde{Q}_3 written in the basis (2.3) is not more than linear in c and Δ ,

$$\tilde{Q}_3 = c\,\tilde{Q}_3^c + \Delta\,\tilde{Q}_3^\Delta + \tilde{Q}_3^{(0)}.\tag{2.30}$$

Matrices \tilde{Q}^c_3 , \tilde{Q}^Δ_3 have lower-triangular form with the diagonal elements being

$$\ell^3 \tilde{Q}_3^c | m_i, \Delta \rangle = \lambda | m_i, \Delta \rangle + \dots, \qquad \lambda = \frac{1}{6} \left(\sum_i m_i^3 - m_i \right),$$
 (2.31)

$$\ell^3 \tilde{Q}_3^{\Delta} | m_i, \Delta \rangle = \nu | m_i, \Delta \rangle + \dots, \qquad \nu = 4 \, n. \tag{2.32}$$

Because of the triangular form, $c \lambda + \Delta \nu$ are the eigenvalues of $c \tilde{Q}_3^c + \Delta \tilde{Q}_3^{\Delta}$. To estimate contribution of each term in (2.30) toward free energy in the appendix A we calculate

$$\ell^3 \langle \tilde{Q}_3 \rangle_{\Delta,n} \equiv \frac{\ell^3}{P(n)} \sum_{\{m\}=n} \langle m_i, \Delta | \tilde{Q}_3 | m_i, \Delta \rangle = (a_0 c + b_0) n^2 + 4\Delta n + O(\ell^3), \qquad (2.33)$$

in the limit of infinite n, assuming $n, \Delta \sim \ell^2$. The n-independent coefficients

$$a_0 = \frac{2}{5}, \qquad b_0 = 4. (2.34)$$

We note that leading n-scaling in (2.33) is fixed to be n^2 lest the contribution of \tilde{Q}_3 diverge in the thermodynamic limit $\ell \to \infty$. At leading order in c this is easy to see directly from the sum over Young tableaux in (2.33),

$$\frac{1}{P(n)} \sum_{\{m\}=n} \left(\frac{1}{6} \sum_{i} m_i^3 \right) = a_0 n^2 + O(n). \tag{2.35}$$

For large n there are exponentially many ways to represent n as a sum of integers with the typical partition consisting of $\sim \sqrt{n}$ terms with each term being of order $m_i \sim \sqrt{n}$. From here it immediately follows that $\sum_i m_i^3 \sim \sqrt{n} \times n^{3/2} \sim n^2$. After substituting scaling of saddle-point values $\Delta^* \sim c\ell^2$, $n^* \sim \ell^2$ into (2.33) we find that both $c \tilde{Q}_3^c$ and $\Delta \tilde{Q}_3^\Delta$ contribute toward f_1 and potentially to f_2 , while $\tilde{Q}_3^{(0)}$ contributes to f_2 only. So far we are interested only in f_1 we can simplify (2.14) to be

$$e^{\tilde{\mathcal{L}}(c,\mu_3/\ell^3,\Delta,n)} \equiv \frac{e^{-\frac{\mu_3}{\ell^3}4n\Delta}}{P(n)} \sum_{\{m\}=n} e^{-\frac{\mu_3}{\ell^3}\frac{c}{6}\sum_i m_i^3},$$
(2.36)

$$\tilde{\mathcal{L}} = -\mu_3/\ell^3 \, 4 \, n \, \Delta - a_0 \, (c \, \mu_3/\ell^3) n^2 - a_1 \, (c \, \mu_3/\ell^3)^2 n^{7/2} - a_2 \, (c \, \mu_3/\ell^3)^3 n^5 + \dots$$
(2.37)

The expansion (2.37) assumes $\ell^3/(c\,\mu_3) \gg n \gg 1$. Sum over Young tableaux (2.36) provides a non-perturbative definition of $\tilde{\mathcal{L}}$. It is a simplified version of the sums appearing in [36]. A few first coefficients a_i are calculated in the appendix B.

In principle effective action (2.37) together with (2.16) completely determines $f_1(t_3)$ via maximization over Δ and n, but there is a direct way to obtain f_1 bypassing this step. We first notice that at leading order in c effective action (2.16) as a function of Δ reduces to (2.19) and therefore saddle-point value of Δ^* is fixed by (2.22) independently of the value of n^* ,

$$\Delta^* = \frac{c\pi^2\ell^2}{6\,\beta^2} \left(e^* + O(1/c)\right). \tag{2.38}$$

In case when only t_3 is "turned on" e^* can be found explicitly,

$$e^* = \frac{\left(t_3 - \left(t_3^3 + 27t_3^4 + 3^{3/2}t_3^{7/2}\sqrt{2 + 27t_3}\right)^{1/3}\right)^2}{6t_3^2 \left(t_3^3 + 27t_3^4 + 3^{3/2}t_3^{7/2}\sqrt{2 + 27t_3}\right)^{1/3}}$$
$$= 1 - 4t_3 + 28t_3^2 - 240t_3^3 + 2288t_3^4 - 23296t_3^5 + O\left(t_3^6\right).$$

Then leading 1/c correction to free energy can be found now by plugging Δ^* back into (2.16) and keeping $O(c^0)$ terms,

$$e^{\frac{\pi^2 \ell}{6\beta} f_1} = e^{-\frac{\pi}{2} \sqrt{\frac{\Delta^*}{6c}}} \sum_{\{m_i\}} e^{-\frac{\beta}{\ell} n - 6\frac{\mu_3}{\ell^3} n\Delta^* - \frac{\mu_3}{\ell^3} \frac{c}{6} \sum_i m_i^3}.$$
 (2.39)

Here $n = \sum_i m_i$ and the sum goes over all Young tableaux. This sum can be calculated by rewriting (2.39) using free boson representation.² Namely, each set m_i will be parametrized by the set of integer numbers r_k , where r_k is the number of times natural number k appears in the set m_i . Then the sum over all sets $\{m_i\}$ is equivalent to the sum over all r_k and

$$\sum_{i} m_{i} = \sum_{k=1}^{\infty} k \, r_{k}, \qquad \sum_{i} m_{i}^{3} = \sum_{k=1}^{\infty} k^{3} r_{k}. \tag{2.40}$$

For any coefficients x, y > 0 we find

$$\sum_{\{m_i\}} e^{-x\sum_i m_i - y\sum_i m_i^3} = \sum_{r_1=0}^{\infty} \sum_{r_2=0}^{\infty} \dots e^{-\sum_k r_k (xk + yk^3)} = \prod_{k=1}^{\infty} \left(1 - e^{-xk - yk^3}\right)^{-1}.$$
 (2.41)

This infinite product can be consequently rewritten as an exponent of the sum of logarithms. Going back to (2.39) this gives

$$\frac{\pi^2 \ell}{6\beta} f_1 = -\frac{\pi^2 \ell}{6\beta} \sqrt{e^*} - \sum_{k=1}^{\infty} \log \left(1 - e^{-\frac{\beta}{\ell} (1 + 6t_3 e^*)k - \frac{t_3}{\pi^2} \left(\frac{\beta}{\ell}\right)^3 k^3} \right). \tag{2.42}$$

The effective variables in (2.42) is the combination $\beta k/\ell$ and in the limit $\ell \to \infty$ the sum over k can be substituted by an integral, yielding

$$f_1 = -\sqrt{e^*} - \frac{6}{\pi^2} \int_0^\infty dk \log\left(1 - e^{-(1 + 6t_3 e^*)k - t_3 k^3/\pi^2}\right). \tag{2.43}$$

A few first terms in Taylor series expansion of f_1 can be easily calculated by expanding this expression in t_3 ,

$$f_1 = -\frac{22}{5}t_3 + \frac{2096}{35}t_3^2 - \frac{4464}{5}t_3^3 + \frac{82304}{5}t_3^4 + O\left(t_3^5\right). \tag{2.44}$$

This agrees with the perturbative calculation of [35].

2.2.2 1/c corrections from \tilde{Q}_5

Adding μ_5 to consideration is straightforward. We notice that \tilde{Q}_5 is not more than quadratic in c and Δ ,

$$\tilde{Q}_{5} = c^{2} \tilde{Q}_{5}^{cc} + c \Delta \tilde{Q}_{5}^{c\Delta} + \Delta^{2} \tilde{Q}_{5}^{\Delta\Delta} + c \tilde{Q}_{5}^{c} + \Delta \tilde{Q}_{5}^{\Delta} + \tilde{Q}_{5}^{(0)}, \tag{2.45}$$

and all three matrices \tilde{Q}^{cc} , $\tilde{Q}^{c\Delta}$, $\tilde{Q}^{\Delta\Delta}_5$ are lower-triangular in the basis (2.3). Their diagonal matrix elements are easy to calculate,

$$\ell^5 \tilde{Q}_5^{cc} | m_i, \Delta \rangle = \alpha | m_i, \Delta \rangle + \dots, \qquad \alpha = \frac{1}{12} \left(\sum_i \frac{m_i^5}{6} - \frac{5m_i^3}{12} - \frac{m_i}{4} \right),$$
 (2.46)

$$\ell^5 \tilde{Q}_5^{c\Delta} | m_i, \Delta \rangle = \delta | m_i, \Delta \rangle + \dots, \qquad \delta = \sum_i \frac{5}{6} m_i^3 - m_i, \tag{2.47}$$

$$\ell^5 \tilde{Q}_5^{\Delta \Delta} | m_i, \Delta \rangle = \gamma | m_i, \Delta \rangle + \dots, \qquad \gamma = 12 \, n.$$
 (2.48)

²We thank Nikita Nekrasov for help with the following calculation.

Only these three matrices contribute toward f_1 via

$$e^{\tilde{\mathcal{L}}(c,\mu_3/\ell^3,\mu_5/\ell^5,\Delta,n)} \equiv \frac{e^{-\frac{\mu_3}{\ell^3}4n\Delta - \frac{\mu_5}{\ell^5}12n\Delta^2}}{P(n)} \sum_{\{m\}=n} e^{-\left(\frac{c\mu_3}{6\ell^3} + \frac{5c\Delta\mu_5}{6\ell^5}\right)\sum_i m_i^3 - \frac{c^2\mu_5}{72\ell^5}\sum_i m_i^5}.$$
(2.49)

Next steps are completely analogous to the case with only Q_3 . At leading order in c saddle point value of Δ is fixed by

$$s_0 = 2\sqrt{e} - e - t_3 e^2 - t_5 e^3, (2.50)$$

$$f_0 = s_0(e^*), \qquad \frac{ds_0}{de}\Big|_{e^*} = 0, \qquad \Delta^* = \frac{c\pi^2\ell^2}{6\,\beta^2}e^*,$$
 (2.51)

while f_1 is given by

$$f_1 = -\sqrt{e^*} - \frac{6}{\pi^2} \int_0^\infty dk \log\left(1 - e^{-(1+6t_3e^* + 15t_5(e^*)^2)k - (t_3 + 5t_5e^*)k^3/\pi^2 - \frac{1}{2}t_5k^5/\pi^4}\right). \tag{2.52}$$

A few first terms in Taylor series expansion are

$$f_1 = -\frac{22}{5}t_3 - \frac{302}{21}t_5 + \frac{2096}{35}t_3^2 + \frac{14328}{35}t_3t_5 + \frac{51168}{77}t_5^2 + \dots$$
 (2.53)

Generalization to include higher charges is conceptually straightforward. The leading term f_0 is given by (2.22) with $s_0 = 2\sqrt{e} - e - \sum_{j=2}^{\infty} t_{2j-1}e^j$. We expect that all higher charges \tilde{Q}_{2j-1} also have lower-triangular form at leading in order c, and therefore (2.29) will reduce to an appropriate sum over Young tableaux, yielding an appropriate generalization of (2.52).

3 Comparison of GGE with eigenstate ensemble

3.1 Eigenstate Thermalization Hypothesis

Eigenstate Thermalization Hypothesis (ETH), developed in the works of Deutsch and Srednicki in the 90s [37, 38] (also see [39, 40] for an earlier work on the subject) is the idea that individual energy eigenstate of a sufficiently complex "chaotic" many-body quantum system exhibits thermal properties. ETH provides a mechanism explaining thermalization of isolated quantum systems. Extensive numerical studies during last decade have supported the expectation that ETH is a general property of quantum many-body systems, unless the system exhibits an extensive number of conserved quantities [41].

At the colloquial level the ETH promotes a highly excited energy eigenstate of a quantum system to the "eigenstate ensemble" stating that the latter can describe thermal properties of a quantum system in the same way as the conventional canonical and micro-canonical ensembles. The discrepancy between e.g. expectation values of local quantities in these different ensembles then would be suppressed in the thermodynamic limit. Holographic CFTs are expected to be complex enough to exhibit thermalization starting from a sufficiently excited pure state. On the dual gravity side this is a process of black hole formation via gravitational collapse. Consequently holographic CFTs are expected to exhibit eigenstate thermalization, at least in some form.

A basic argument of [42] shows that CFTs can not satisfy the conventional form of ETH when all sufficiently excited energy eigenstates are thermal. This is because in the latter case all eigenstates would have to exhibit the same thermal properties: expectation value of a local operator \mathcal{O} would be a function of effective temperature, i.e. would depend on energy only

$$\langle E|\mathcal{O}|E\rangle = \mathcal{O}_{\text{eth}}(\beta(E)).$$
 (3.1)

Operator-state correspondence allows reformulation (3.1) in terms of OPE coefficients. Then it can be easily shown that the primary states and the descendants can not both satisfy (3.1) with the same function \mathcal{O}_{eth} .

It was proposed in [43] that the non-trivial content of ETH is not the equivalence between individual eigenstates and thermal ensemble, but the equivalence of individual eigenstates from a certain class with each other. Then the equivalence with the thermal ensemble would follow automatically, provided the original class of eigenstates is wide enough. This idea lead [42] to propose the following formulation of local ETH in CFTs: (3.1) should apply to any local operator \mathcal{O} but be limited only to primary (Virasoro-primary in 2d) states $|E\rangle$. In this formulation $\mathcal{O}_{\text{eth}}(\beta)$ is not necessarily related to thermal expectation value of \mathcal{O} . For conformal theories in $d \geq 3$ it was further argued in [44] that the primary states dominate the microcanonical ensemble and therefore in the thermodynamic limit $\mathcal{O}_{\text{eth}}(\beta)$ coincides with the conventional thermal expectation value of \mathcal{O} . The d=2 case is a subject of discussion below.

In 2d we need to distinguish two cases: when \mathcal{O} is a Virasoro descendant of identity or not. In the former case the local ETH (3.1) is automatic — the corresponding heavy-heavy-light OPE coefficient is fixed by Virasoro algebra and is a smooth function of E, dimension of the Virasoro primary state $|E\rangle$. In the latter case, when \mathcal{O} is not a descendant of identity, there are no known explicit examples when (3.1) is satisfied with a smooth $\mathcal{O}_{\text{eth}}(\beta)$. There is a general expectation though that this is the case for certain large central charge theories, including holographic CFTs. In d=2 thermal expectation value of any such \mathcal{O} is zero (because thermal cylinder is conformally flat). It is therefore often assumed that in sufficiently complex 2d theories corresponding heavy-heavy-light OPE is suppressed by the dimension of the heavy operator. Dominance of the vacuum conformal family or the "identity block" in the OPE of two heavy primaries is an underlying assumption in many works on large central charge theories in the context of ETH and thermalization [10, 14, 42, 45–53].

Leaving aside the behavior of \mathcal{O} outside of vacuum conformal family, below we focus on the case when \mathcal{O} is a Virasoro descendant of identity. Local ETH (3.1) is automatic in this case, but function $\mathcal{O}_{eth}(\beta)$ a priory has no interpretation in terms of thermal physics. A natural question then would be to compare $\mathcal{O}_{eth}(\beta)$ with thermal expectation value of the operator \mathcal{O} ,

$$\mathcal{O}_{\rm th}(\beta) = \frac{1}{Z(\beta)} \text{Tr}(\mathcal{O} e^{-\beta H}).$$
 (3.2)

It was expected that at infinite central charge locally eigenstate is equivalent to the thermal ensemble [6], which is indeed the case, $\mathcal{O}_{\mathrm{eth}} = \mathcal{O}_{\mathrm{th}}$. At the same time $\mathcal{O}_{\mathrm{eth}}$ and $\mathcal{O}_{\mathrm{th}}$ do not

match at the subleading order in 1/c [44, 46–49]. A naive explanation of this discrepancy is that the "eigenstate ensemble" $|E\rangle$ has positive value of qKdV charges and hence should be compared not with (3.2) but with the full Generalized Gibbs Ensemble

$$\mathcal{O}_{GGE}(\beta, \mu_i) = \frac{1}{Z(\beta, \mu_i)} \text{Tr}\left(\mathcal{O} e^{-\beta H - \sum_k \mu_{2k-1} Q_{2k-1}}\right), \tag{3.3}$$

with the fugacities μ_i chosen to match quantum numbers of $|E\rangle$. This is the comparison performed below.

3.2 Comparison at infinite central charge

From now on we restrict to the case when $|E\rangle$ is a heavy scalar Virasoro primary. To achieve finite energy density and thus finite effective temperature in the thermodynamic limit the dimension E should scale as ℓ^2 ,

$$\ell \to \infty, \qquad E/\ell^2 = \text{fixed.}$$
 (3.4)

A Virasoro descendant of identity \mathcal{O} can be either a quasi-primary or a total derivative. In the latter case by Lorentz invariance $\mathcal{O}_{\text{eth}} = \mathcal{O}_{\text{GGE}} = 0$. When \mathcal{O} is a quasiprimary of dimension 2k, \mathcal{O}_{eth} is non-trivial in the thermodynamic limit (3.4) only if \mathcal{O} includes $\underbrace{(T\ldots(TT))}_{\text{k times}}$. One can always choose a basis at the level 2k such that a unique quasi-

primary with non-vanishing expectations value (3.1) in the thermodynamic limit is the density of KdV charge $\mathcal{O} = q_{2k-1}$ [44],

$$Q_{2k-1} \equiv \int_0^\ell du \, q_{2k-1}. \tag{3.5}$$

To investigate the equality of the eigenstate and generalized Gibbs ensembles we compare

$$\langle E|q_{2k-1}|E\rangle = \frac{1}{\ell}\langle E|Q_{2k-1}|E\rangle = \left(\frac{E}{\ell^2}\right)^k, \qquad E,\ell \to +\infty,$$
 (3.6)

and

$$\langle q_{2k-1}\rangle_{\text{GGE}} = \frac{1}{\ell} \langle Q_{2k-1}\rangle_{\text{GGE}} \equiv -\frac{1}{\ell} \partial_{\mu_{2k-1}} \log Z,$$
 (3.7)

for $k \geq 2$. For k = 1 the equality between $\langle E|T|E\rangle$ and $\langle T\rangle_{GGE}$,

$$E = \langle H \rangle_{\text{GGE}} \tag{3.8}$$

is the relation which defines effective temperature β in terms of E/ℓ^2 . The expectation values (3.6) satisfy $\langle E|q_{2k-1}|E\rangle = \langle E|q_1|E\rangle^k$, $q_1 \equiv T$. Therefore for the equality between the eigenstate and generalized Gibbs ensembles to hold it is necessary for the GGE partition function to satisfy (2.24). This is the case at leading order in central charge for any values of μ_{2k-1} as discussed in section 2.1. Hence we establish that for the descendant of identity $\mathcal{O}=q_{2k-1}$, at leading order in central charge expectation value in a primary state is the same as in the generalized Gibbs ensemble for any choice of μ_{2k-1} . The same conclusion,

about the equivalence of the eigenstate and the GGE for any choice of μ_{2k-1} also trivially applies to all other descendants of identity, because for them at leading order in central charge both (3.1) and (3.3) are zero. Equivalence for all μ_{2k-1} is consistent with the holographic interpretation that at infinite central charge classical black hole in AdS₃ is dual to (1.5) for any values of μ_{2k-1} [29]. Also see [54] for bulk interpretation of the KdV charges in terms of boundary gravitons.

3.3 Discrepancy at finite central charge

At finite central charge the relations (2.24) are not automatically satisfied. Using the expansion (2.28), at leading order in 1/c one finds from (2.25), (2.26),

$$2(f_0 + 3t_3\partial_{t_3}f_0 + 5t_5\partial_{t_5}f_0 + \dots)(f_1 + 3t_3\partial_{t_3}f_1 + 5t_5\partial_{t_5}f_1 + \dots) + \partial_{t_3}f_1 = 0,$$
 (3.9)

$$3(f_0 + 3t_3\partial_{t_3}f_0 + 5t_5\partial_{t_5}f_0 + \dots)^2(f_1 + 3t_3\partial_{t_3}f_1 + 5t_5\partial_{t_5}f_1 + \dots) + \partial_{t_5}f_1 = 0, \quad (3.10)$$

. . .

In principle one can hope these equations would specify a set of t_{2k-1} such that (3.9), (3.10), ... are satisfied, which would ensure equivalence between the eigenstate and generalized Gibbs ensembles at the level of the expectation values of q_{2k-1} at first subleading order in 1/c. It should be noted though that the equations (3.9), (3.10), ... have no small parameter, because when written in terms of variables t_{2k-1} they are c-independent. Thus one would need to know full function $f_1(t_3, t_5, ...)$ to find a possible solution. In this sense the problem of matching fugacities μ_{2k-1} to a primary state is non-pertrubative, i.e. it requires knowledge of free energy at all orders in μ_{2k-1} , even for large c [44].

Although we do not know the full function $f_1(t_3, t_5, ...)$, a simple argument shows that a solution of (3.9) and more generally of (2.24) does not exist for finite but large c. Let us compare (3.6) with (3.7) for Q_3 by calculating the difference between the two,

$$\ell^{-1}\left(\langle Q_3\rangle_{\text{GGE}} - \langle E|Q_3|E\rangle\right) = \ell^{-1}\left(\langle \hat{Q}_3\rangle_{\text{GGE}} + \langle \tilde{Q}_3\rangle_{\text{GGE}}\right) - \left(\frac{E}{\ell^2}\right)^2. \tag{3.11}$$

Using explicit expression for \hat{Q}_3 (2.11) we find that in the thermodynamic limit the expectation value of \hat{Q}_3 is given by the saddle point,

$$\ell^{-1}\langle \hat{Q}_3 \rangle_{\text{GGE}} = \left(\frac{E^*}{\ell^2}\right)^2 + O(1/\ell). \tag{3.12}$$

The saddle point value $E^* = \Delta^* + n^*$ is equal to the energy E of state $|E\rangle$ due to (3.8). The values of Δ^* , n^* should be determined from the full effective action $\mathcal{L} + \tilde{\mathcal{L}}$ (2.16), (2.29). The full effective action is not known, but both Δ^* and n^* scale with the system size as $\sim \ell^2$. Furthermore, up to 1/c corrections $\Delta^* \approx E^* = E$ and therefore Δ^*/ℓ^2 is a positive number in the thermodynamic limit. Using (3.11) and (3.12) we find the discrepancy between the eigenstate and GGE expectation values to be

$$\ell^{-1} \left(\langle Q_3 \rangle_{\text{GGE}} - \langle E | Q_3 | E \rangle \right) = \ell^{-1} \langle \tilde{Q}_3 \rangle_{\text{GGE}}$$

$$= \frac{e^{-\tilde{\mathcal{L}}(\Delta^*, n^*)}}{\ell P(n^*)} \sum_{\{m\} = n^*} \langle m_i, \Delta^* | \tilde{Q}_3 e^{-\mu_3 \tilde{Q}_3 - \mu_5 \tilde{Q}_5 - \dots} | m_i, \Delta^* \rangle.$$
(3.13)

So far we are interested in the leading in c behavior of (3.13) we can substitute \tilde{Q}_3 by a lower-triangular matrix $c\,\tilde{Q}_3^c + \Delta^*\tilde{Q}_3^\Delta$, \tilde{Q}_5 can be substituted by the lower-triangular $c^2\tilde{Q}_5^{cc} + c\,\Delta^*\tilde{Q}_5^{c\Delta} + (\Delta^*)^2\tilde{Q}_3^{\Delta\Delta}$, and so on. Then average of $\Delta^*\tilde{Q}_3^\Delta$ is simply equal to $4\Delta^*n^*/\ell^3$, while the average of $c\,\tilde{Q}_3^c$ can be rewritten as an average over Young tableaux,

$$\langle \tilde{Q}_{3}^{c} \rangle_{\text{GGE}} = \frac{e^{-\tilde{\mathcal{L}}(\Delta^{*}, n^{*})}}{6 \, \ell^{3} P(n^{*})} \sum_{\{m\}=n^{*}} \left(\sum_{i} m_{i}^{3} \right) e^{-\frac{\mu_{3}}{\ell^{3}} \left(\frac{c}{6} \sum_{i} m_{i}^{3} + 4\Delta^{*} n^{*} \right) - \frac{\mu_{5}}{\ell^{5}} \left(\frac{c^{2}}{72} \sum_{i} m_{i}^{5} + \dots \right) - \dots}{6 \, \ell^{3} P(n^{*})} + O(c^{0}).$$

In the thermodynamic limit this is equal to $a(n^*)^2/\ell^3$, where a is some non-negative function of μ_3, μ_5, \ldots . When all $\mu_{2k-1} = 0$, it reduces to $a_0 = 2/5$, (2.34). Thus we finally have

$$\ell^{-1}(\langle Q_3 \rangle_{GGE} - \langle E|Q_3|E\rangle) = \frac{a(n^*)^2 + 4\Delta^* n^*}{\ell^4} + O(c^0).$$
 (3.14)

The important point here is that the discrepancy (3.14) can be zero if and only if $n^* = 0$. The value of n^* can be interpreted as the effective level of Virasoro descendants which give main contribution to the partition function. It is a priory expected that the discrepancy between the primary state and the GGE would vanish if $n^* = 0$, i.e. if primaries dominate the partition sum. The non-trivial result here is that (3.14) vanishes only if $n^* = 0$.

It is easy to see that primaries do not dominate the partition sum because of the factor P(n) accounting for the exponential growth of the number of descendants with n. Using (2.16), (2.36), (2.37), (2.49) the effective action for n can be rewritten in terms of the variable $n = n/\ell^2$ which remains finite in the thermodynamic limit,

$$\mathcal{L}_{\text{eff}}(\mathbf{n}) = \ell \left(\pi \sqrt{\frac{2\mathbf{n}}{3}} - L \right),$$

$$L = (\beta + 6\mu_3 \Delta^* / \ell^2 + \dots) \mathbf{n} + \mu_3 (1 + ca_0 + b_0 + \dots) \mathbf{n}^2 + O(\mathbf{n}^{5/2}).$$
(3.15)

Since L(n) admits an expansion in powers of n starting from one, n = 0 can not be a solution of $\partial \mathcal{L}_{\text{eff}}/\partial n = 0$ and the maximum of \mathcal{L}_{eff} is achieved at positive $n = n^*/\ell^2$. From here it follows that for large c both $\langle Q_3 \rangle_{\text{GGE}}$ and $\langle E|Q_3|E \rangle$ scale as c^2 , while their difference is non-zero and scales as c. This proves (3.6) and (3.7) for Q_3 are always different for large but finite c for any values of μ_{2k-1} . A similar argument would also apply to Q_5 and higher charges.

4 Discussion

In this paper, we studied Generalized Gibbs Ensemble (1.5) of 2d CFTs in the limit when the size of the spatial circle ℓ goes to infinity. We have observed that qKdV charges Q_{2k-1} can be split into a sum of two terms: first term \hat{Q}_{2k-1} survives in the infinite c limit and is a function of L_0 , while second term \tilde{Q}_{2k-1} is 1/c suppressed and is positive semi-definite. Furthermore leading in 1/c expansion part of \tilde{Q}_{2k-1} assumes a simple lower-triangular form, if written in the conventional basis (2.3). Using this result we have explicitly calculated extensive contribution to free energy in the infinite central charge limit (2.20), (2.22), (2.23), and reduced calculation of leading 1/c correction to a sum over Young tabluex. The latter

can be computed explicitly, provided the expression for the qKdV charge in terms of Virasoro algebra generators is known. We have calculated this correction for the first two qKdV charges Q_3, Q_5 (2.52). Our results suggest adding higher qKdV charges at leading 1/c order will be straightforward, and we hope to report $f_1(t_3, t_5, t_7, ...)$ which includes higher fugacities soon. Since the first two orders in 1/c expansion turned out to be easily tractable, next natural question would be to understand the structure of the GGE partition function at the $1/c^2$ order. We leave this task for the future. Another important direction is to calculate sub-extensive corrections to free energy, even at leading order in c. The logarithmic $\ln \ell$ and $O(\ell^0)$ terms would follow from the saddle point approximation of (2.16), while higher order terms in $1/\ell$ expansion will require more work.

Leading 1/c corrections to the GGE free energy calculated in this paper are of interest from the point of view of AdS/CFT correspondence. They encode quantum corrections on the gravity side and contain information about individual black hole microstates [55]. While from the point of view of local probes the heavy states are thermal at infinite c limit [6, 56], this is no longer the case at finite c. In this paper we compared a heavy Virasoro primary state with the GGE and found that the two are always locally distinguishable no matter what the values of the qKdV fugacities are. At the technical level this comes from the split of Q_3 into \hat{Q}_3 and \tilde{Q}_3 . The expectation value of \hat{Q}_3 is the same in the primary eigenstate and the GGE for any values of the fugacities, but the expectation value of \tilde{Q}_3 is different. The operator \hat{Q}_3 annihilates the primary state, while its expectation value in the GGE at leading order in c is strictly positive. A similar result applies to higher charges as well. This mismatch between the eigenstate and the GGE is surprising. Large central charge theories are expected to be chaotic and thermalize [57], in full consistency with holographic expectations. Yet, we find that in large c CFTs, not only in the non-interacting examples studied previously [27, 58, 59], there are initial configurations labeled by the macroscopic (extensive) values of qKdV charges which fail to thermalize i.e. be described by the GGE upon equilibration.

The breakdown of ergodicity observed in this paper emphasizes the question of content and meaning of the Eigenstate Thermalization Hypothesis in 2d CFTs. First, it should be noted that ETH is always satisfied on average, i.e. when (3.1) is averaged over exponentially many eigenstates with similar energy. In fact the same applies to the off-diagonal matrix elements, on average they are exponentially small and described by a smooth form-factor. These is true for any theory, or in fact any system with a sufficiently large Hilbert space, and can be seen as a consequence of typicality. Correctness of ETH on average was confirmed for 2d CFT recently in [51–53, 60]. As expected the obtained results are valid for any value of central charge, which confirms its kinematic Hamiltonian-independent origin. Arguably, the non-trivial content of ETH should be stronger and apply only to a subclass of chaotic theories. Second, the conventional understanding of Eigenstate Thermalization suggests that (certain) individual energy eigenstates should be compared directly with a thermal ensemble or its kin. We would like to argue for 2d CFTs this is not the correct approach. As was discussed in section 3.1, primary energy eigenstates are indeed thermal in the sense that (3.1) is equal to (3.2), but only at infinite c, while this equality immediately breaks already at first subleaiding 1/c order [44, 46–49]. A natural generalization of this idea

would be to require the primaries to match the GGE. It was shown in this paper to fail at finite c as well. Besides the discrepancies at the level of local observables at finite c, there are other indications that the conventional form of ETH does not apply to primary states. As a non-local probe of thermal behavior one can consider Renyi entropy of an interval, when the size of the interval scales with ℓ . In this limit eigenstate and thermal Renyi entropies are different even when ETH is obeyed, but in the latter case the expression for the eigenstate entropy was conjectured in [43] in terms of the density of states. This conjecture was later studied and verified numerically in [61] for non-integrable spin-chains. At the same time this conjectural form fails badly for large c conformal theories [14]. Another potentially related result is the breakdown of the KMS conditions by a heavy primary state [9]. All this strongly suggests that matching individual primary sates directly to a thermal or generalized thermal ensemble is incorrect. Maybe then the focus should be on the descendants? A simple qualitative argument shows that heavy descendants also can not be thermal. If the state has energy $E = \Delta + n$, $n \gg \Delta$, through Virasoro algebra its properties are completely fixed in terms of $|\Delta\rangle$, which at least at infinite c, matches thermal state at wrong temperature. Non-thermal behavior of heavy descendant states was recently rigorously shown in [50].

To conclude, we suggest that the nontrivial content of ETH in 2d CFTs is that for heavy primary states function \mathcal{O}_{eth} (3.1) is a smooth function of energy, without it being directly related to thermal physics [42]. In this sense our findings do not signal violation of ETH, in fact, we have explained in section 3.1 that for local probes from the vacuum conformal family this version of ETH is trivially satisfied. For this version of ETH to be physically interesting it should be established also for the probes outside of the vacuum family by demonstrating that at least in certain large c theories all heavy-heavy-light OPE coefficients are smooth function of the heavy operator's dimension, or small. This is a challenging task without an immediate ways to approach it. Our point here is that answering this question is crucial to understand if 2d CFTs exhibit an additional structure of heavy energy eigenstates beyond those imposed by symmetries and typicality.

A Computation of a_0 , b_0

In this section we show how to calculate

$$\ell^3 \langle \tilde{Q}_3 \rangle_{\Delta,n} \equiv \frac{\ell^3}{P(n)} \sum_{\{m\}=n} \langle m_i, \Delta | \tilde{Q}_3 | m_i, \Delta \rangle = (a_0 c + b_0) n^2 + 4\Delta n + O(\ell^3), \tag{A.1}$$

in the limit when $\Delta, n \sim \ell^2$. An analogous calculation can be also found in [32]. Recall that $\tilde{Q}_3 = 2 \sum L_{-k} L_k$. The expectation value $\langle \tilde{Q}_3 \rangle_{\Delta,n}$ can be thought of as trace $q^{-n-\Delta} \operatorname{Tr}(q^{L_0} \tilde{Q}_3)$ over a subspace with fixed n and Δ spanned by the states (2.3). Using cyclic property of trace one can easily get [62]

$$\langle L_{-k}L_k\rangle_{\Delta,n} = q^k \langle L_kL_{-k}\rangle_{\Delta,n} = \frac{q^k}{1 - q^k} \langle [L_k, L_{-k}]\rangle_{\Delta,n} =$$

$$= \frac{q^k}{1 - q^k} \left(2k \langle L_0\rangle_{\Delta,n} + \frac{c}{12} (k^3 - k) \right),$$
(A.2)

where $q = e^{-\frac{\beta}{\ell}}$. Summing (A.2) over k one obtains

$$\langle \tilde{Q}_3 \rangle_{\Delta,n} = 4\sigma_1 \langle L_0 \rangle_{\Delta,n} + \frac{c}{6} (\sigma_3 - \sigma_1),$$
 (A.3)

where

$$\sigma_1 = \sum_k \frac{kq^k}{1 - q^k},\tag{A.4}$$

$$\sigma_3 = \sum_k \frac{k^3 q^k}{1 - q^k}.\tag{A.5}$$

Expectation value $\langle L_0 \rangle_{\Delta,n}$ is just equal to $n + \Delta$ and the sums σ_1 , σ_3 can be evaluated in thermodynamic limit by replacing the sum over k by an integral $\sum_k \to \int dk$. In this limit

$$\sigma_1 \to \frac{\ell^2 \pi^2}{6\beta^2},\tag{A.6}$$

$$\sigma_3 \to \frac{\ell^4 \pi^4}{15\beta^4}.\tag{A.7}$$

The final step is to recall that if we sum $\text{Tr}(q^{L_0}\tilde{Q}_3)$ over n main contribution will be given by a particular saddle point value $n=\frac{\ell^2\pi^2}{6\beta^2}$. Thus substituting $\frac{\ell^2\pi^2}{6\beta^2}$ by n we get

$$\langle \tilde{Q}_3 \rangle_{\Delta,n} = 4n(n+\Delta) + \frac{2}{5}c n^2.$$
 (A.8)

As a result, $a_0 = \frac{2}{5}$ and $b_0 = 4$. It is straightforward but tedious to generalize the above calculation to higher orders to obtain a_1 , a_2 etc. from (2.37).

B Effective action at 1/c

In this section we show how to calculate the effective action (2.37), which is a partition function decorated by the first non-trivial qKdV charge Q_3 restricted to a particular large descendant level n,

$$e^{\tilde{\mathcal{L}}(c,\mu_3/\ell^3,\Delta,n)} \equiv \frac{e^{-\frac{\mu_3}{\ell^3}4n\Delta}}{P(n)} \sum_{\{m\}=n} e^{-\frac{\mu_3}{\ell^3}\frac{c}{6}\sum_i m_i^3}.$$
 (B.1)

We remind the reader that first 1/c correction to free energy f_1 calculated in section 2.2.1 is given by (compare with (2.39))

$$e^{\frac{\pi^2 \ell}{6\beta} f_1} = e^{-\frac{\pi}{2} \sqrt{\frac{\Delta^*}{6c}}} \sum_n P(n) e^{-\frac{\beta}{\ell} n - \frac{\mu_3}{\ell^3} 2n\Delta^* + \tilde{\mathcal{L}}(c, \mu_3/\ell^3, \Delta^*, n)}.$$
 (B.2)

To calculate $\tilde{\mathcal{L}}$ we introduce an auxiliary partition function of the same kind, which depends on arbitrary parameter x,

$$e^{F(x)} = \sum_{n} P(n) e^{-\frac{x}{\ell}n + \frac{\mu_3}{\ell^3} 4n\Delta + \tilde{\mathcal{L}}(c, \mu_3/\ell^3, \Delta, n)}.$$
 (B.3)

In the limit of large ℓ this sum is saturated at some saddle point n^* ,

$$F(x) = \pi \sqrt{\frac{2n^*}{3}} - \frac{x}{\ell} n^* + \frac{\mu_3}{\ell^3} 4n^* \Delta + \tilde{\mathcal{L}}(c, \, \mu_3/\ell^3, \Delta, n^*), \tag{B.4}$$

which is related to x and μ_3 via $n^* = -\ell \frac{dF}{dx}$. "Free energy" F can be calculated exactly the same way f_1 was calculated in section 2.2.1,

$$F(x) = -\ell \int_0^\infty dk \ln \left(1 - e^{-xk - \frac{c}{6}\mu_3 k^3} \right).$$
 (B.5)

After changing the integration variable $k \to k/x$ and expanding in μ_3 one finds

$$F(x) = \frac{\ell}{x} \left(\frac{\pi^2}{6} - \frac{\pi^4 c \mu_3}{90x^3} + \frac{2\pi^6 c^2 \mu_3^2}{189x^6} - \frac{4\pi^8 c^3 \mu_3^3}{135x^9} + \frac{40\pi^{10} c^4 \mu_3^4}{243x^{12}} + \dots \right)$$
(B.6)

Using the explicit expression (B.5) we can relate n^* and x,

$$\frac{n^*}{\ell^2} = \frac{d}{dx} \int_0^\infty dk \ln\left(1 - e^{-xk - \frac{c}{6}\mu_3 k^3}\right),\tag{B.7}$$

and solve this equation perturbatively with respect to x by expanding in μ_3 ,

$$x = \frac{\pi \ell}{\sqrt{6n^*}} - \frac{4c\mu_3 n^*}{5\ell^2} + \frac{144\sqrt{6}c^2\mu_3^2(n^*)^{5/2}}{25\pi\ell^5} - \frac{20736c^3\mu_3^3(n^*)^4}{25\pi^2\ell^8} + \frac{4762368\sqrt{6}c^4\mu_3^4(n^*)^{11/2}}{125\pi^3\ell^{11}} + \dots$$

Plugging this back into (B.4) and finally renaming n^* into n gives (compare with (2.37))

$$\frac{\mu_3}{\ell^3} 4n\Delta + \tilde{\mathcal{L}}(c, \mu_3/\ell^3, \Delta, n) = -\pi \sqrt{\frac{2n}{3}} + \frac{x(n)}{\ell} n + F(x(n))$$

$$= -\frac{2}{5} (c \,\mu_3/\ell^3) n^2 + \frac{288\sqrt{6}}{175\pi} (c \,\mu_3/\ell^3)^2 n^{7/2} - \frac{20736}{125\pi^2} (c \,\mu_3/\ell^3)^3 n^5$$

$$+ \frac{732672\sqrt{6}}{125\pi^3} (c \,\mu_3/\ell^3)^4 n^{13/2} \dots$$
(B.9)

Generalization to include Q_5 and higher charges would be tedious but straightforward.

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